

# Ultra-High Vacuum Thermal and Carbothermal Reduction of Lunar Regolith Simulants for Oxygen and Construction Materials

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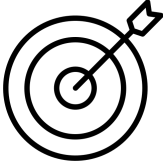

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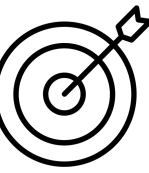
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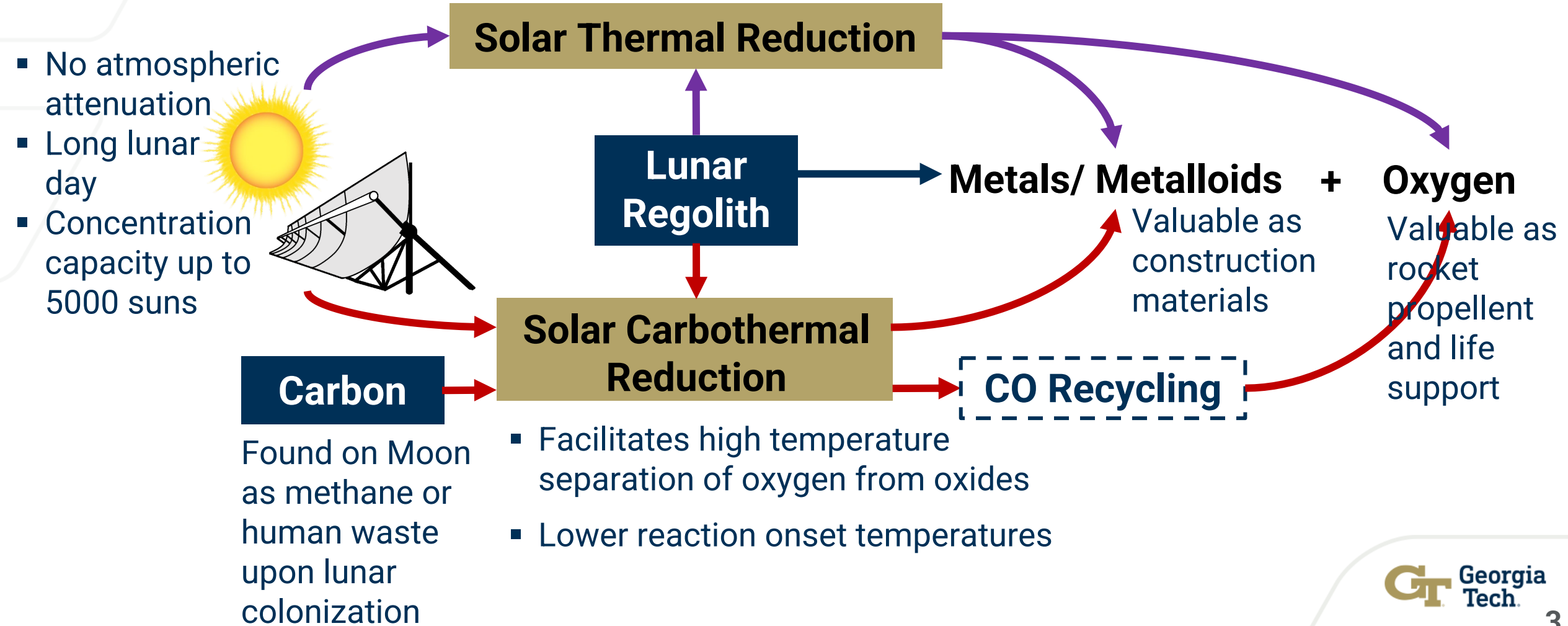
# Outline

- Background and Motivation 
- Methodologies 
- Results 
  - Material characterization 
- Summary and conclusions 

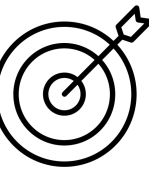
# Background and Motivation



Thermochemical processing of lunar regolith to produce oxygen and construction materials for ISRU to advance space exploration efforts



# Lunar Regolith Simulants



JSC-1A



- Simulating Apollo sample 14163
- Oxides (weight %):
  - $\text{SiO}_2$  – 47.4%
  - $\text{Al}_2\text{O}_3$  – 16.1%
  - $\text{Fe}_2\text{O}_3$  – 11.4%

LMS-1



- Generic mare
- Oxides (weight %):
  - $\text{SiO}_2$  – 46.9%
  - $\text{MgO}$  – 16.8%
  - $\text{Al}_2\text{O}_3$  – 12.4%

LHS-1

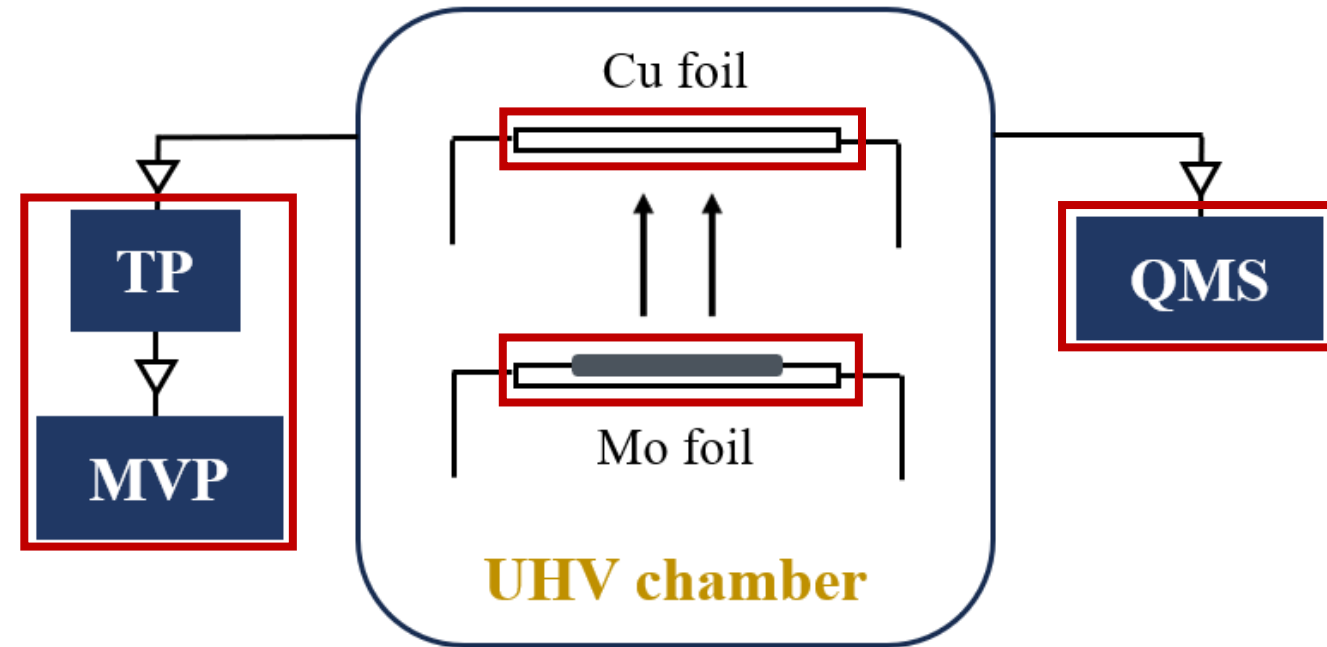


- Generic highland
- Oxides (weight %):
  - $\text{SiO}_2$  – 51.2%
  - $\text{Al}_2\text{O}_3$  – 26.6%
  - $\text{CaO}$  – 12.8%

# UHV Experiment Methodology



- Continuously evacuated via turbomolecular in series with mechanical vacuum pump to maintain  $\sim 10^{-12}$  bar, imitating low-pressure lunar surface conditions
- Sample of  $\sim 5$  mg spread on an Mo foil, and Cu foil placed  $\sim 3$  mm above the sample for vapor depositions
- Quadrupole mass spectrometry (QMS) attached to UHV chamber for measuring gas evolution ( $O_2$  or  $CO$ )





# UHV Experiment Methodology: Direct Thermal

- Two experiments conducted consecutively using the same sample for each simulant
- Thermodynamic analyses conducted at  $\sim 10^{-12}$  bar to determine theoretical operating temperatures for targeted metal production
- Si(g) was not targeted due to requirement of  $T \geq 2000$  °C for production
- SiO(g) was targeted instead as it is an intermediary in Si production
- Isotherm duration was shortened at higher  $P$  runs to avoid Mo sample holder deterioration

Simulant	$P, W$	$T, ^\circ C$		Targeted metal	Isotherm duration, min
		Target	Measured		
LMS-1	613.2	1170	1182-1195	Mg(g), SiO(g)	60
	730.0	1380	1234-1245	Al(g), SiO(g)	30
LHS-1	835.9	1300	1304-1313	Ca(g), SiO(g)	30
	985.5	1500	1484-1498	Al(g), SiO(g)	5





# UHV Experiment Methodology: Carbothermal

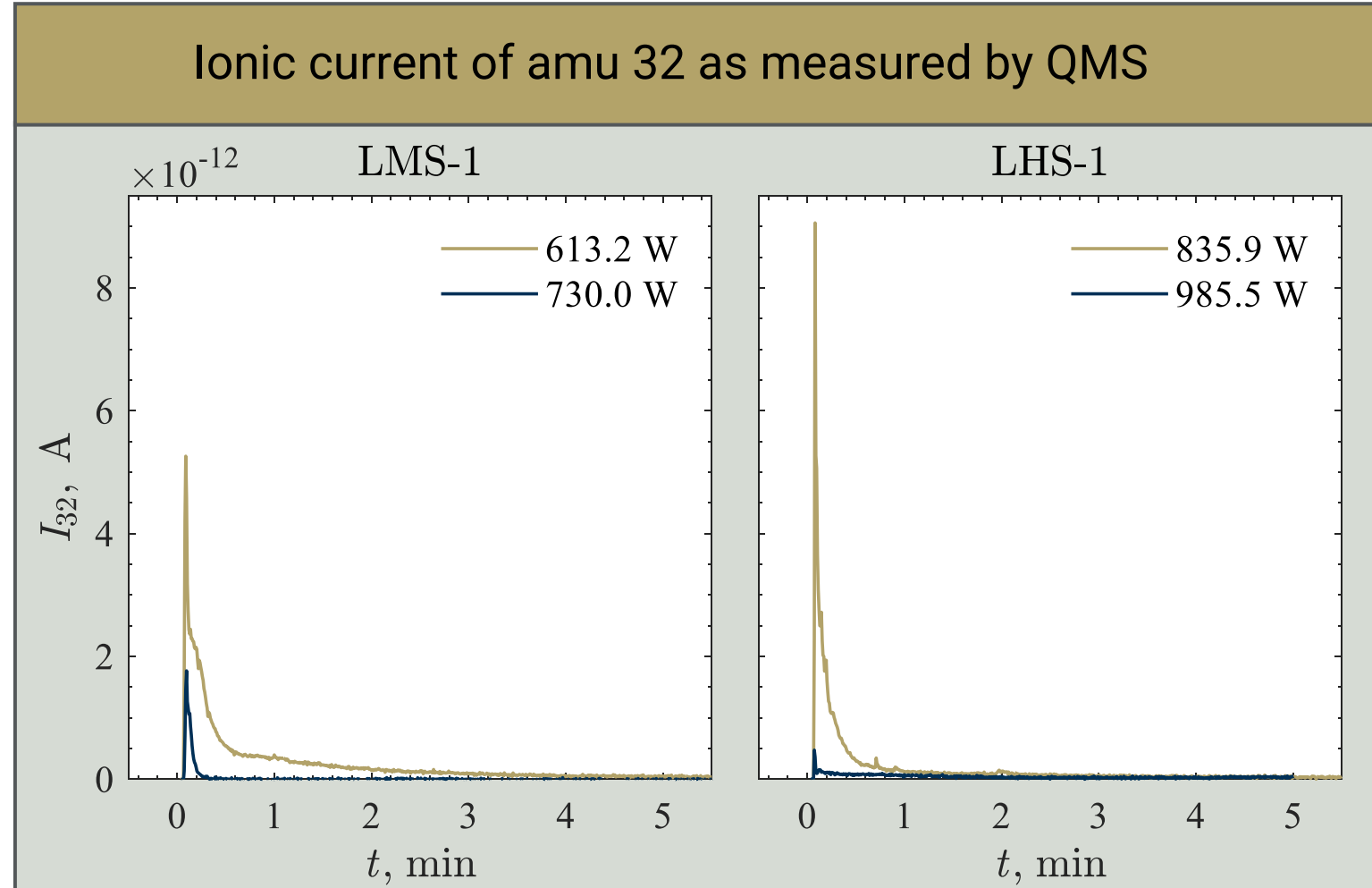
- Simulants were mechanically mixed with pyrolyzed activated carbon at C to simulant mass ratio of 0.5. This ratio represents larger amounts of C than theoretically needed to reduce all oxides in the simulants.
- Two experiments conducted consecutively using the same sample for each simulant mixture
- $T \geq 1200\text{ }^{\circ}\text{C}$  and  $\geq 1400\text{ }^{\circ}\text{C}$  were chosen since previous study has shown that  $T > 1200\text{ }^{\circ}\text{C}$  is needed for vapor production

Simulant mixture	$P, W$	$T, ^{\circ}\text{C}$		Isotherm duration, min
		Target	Measured	
JSC-1A + C	766.5	$\geq 1200$	1243-1256	60
	912.5	$\geq 1400$	1415-1432	30
LHS-1 + C	803.0	$\geq 1200$	1366-1380	60
	876.0	$\geq 1400$	1455-1458	30



# Direct Thermal Reduction

- $O_2(g)$  evolution from sample during experiments was evidenced by the QMS at 32 amu
- A prominent peak at  $t \leq 0.5$  min was observed with no significant changes in  $I_{32}$  for  $t > 4$  min, indicative of rapid thermal reductions
- Smaller amounts of  $O_2(g)$  evolved during the second experiments because the reacted sample from first experiments were used causing smaller sample size



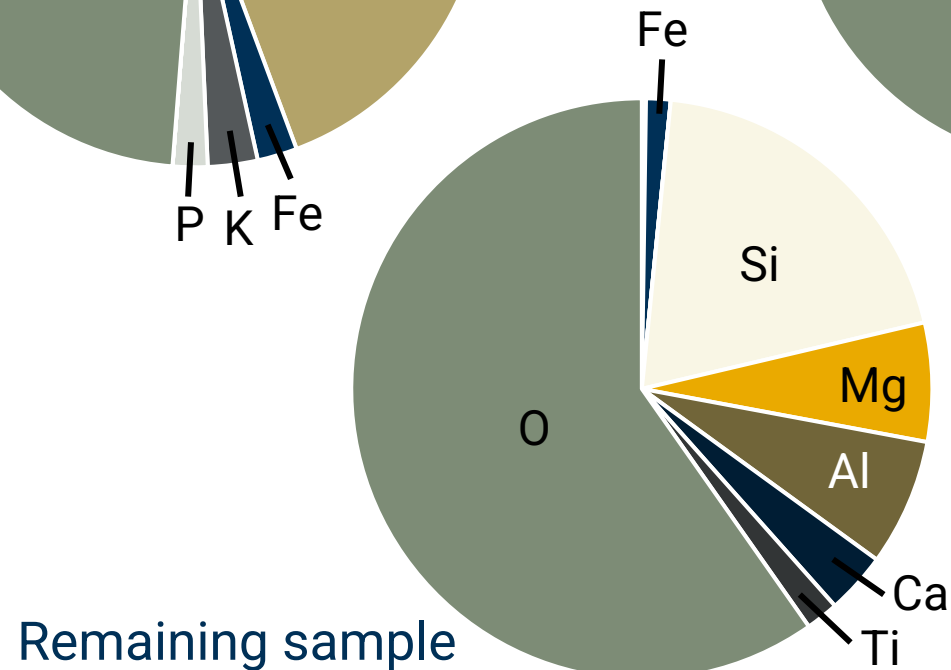
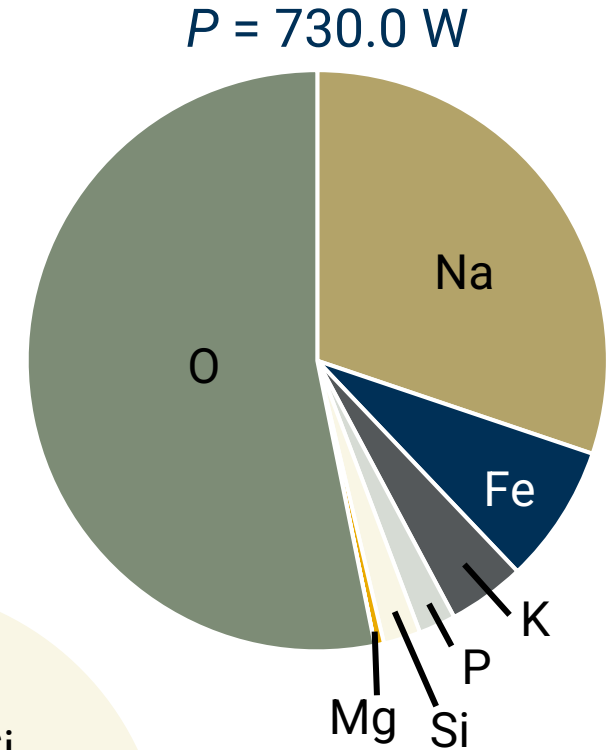
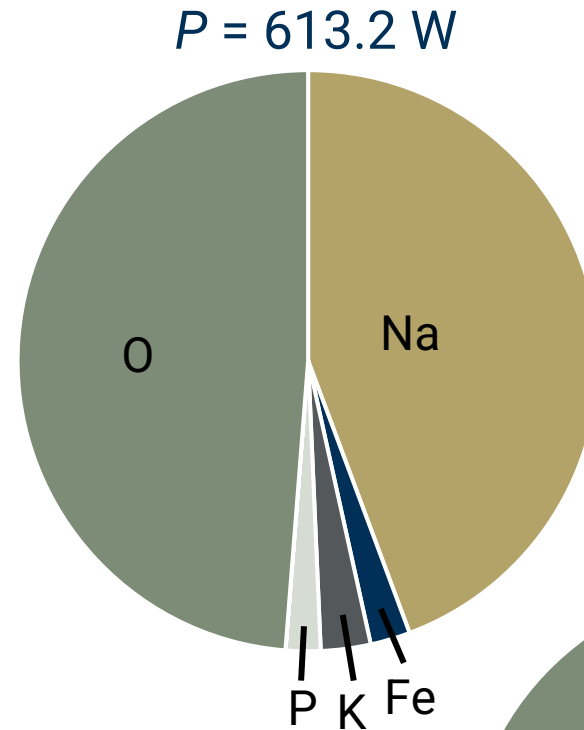


# Direct Thermal Reduction: Material Characterization



## LMS-1 – EDS

- Depositions were evenly distributed on all foils
- Na, Fe, K, and P evolved during both experiments suggesting separation possibility at reasonable temperatures
- Si and Mg evolved during the second experiment and Al, Ca, and Ti remained in sample upon experimentation evidencing requirement of elevated temperatures

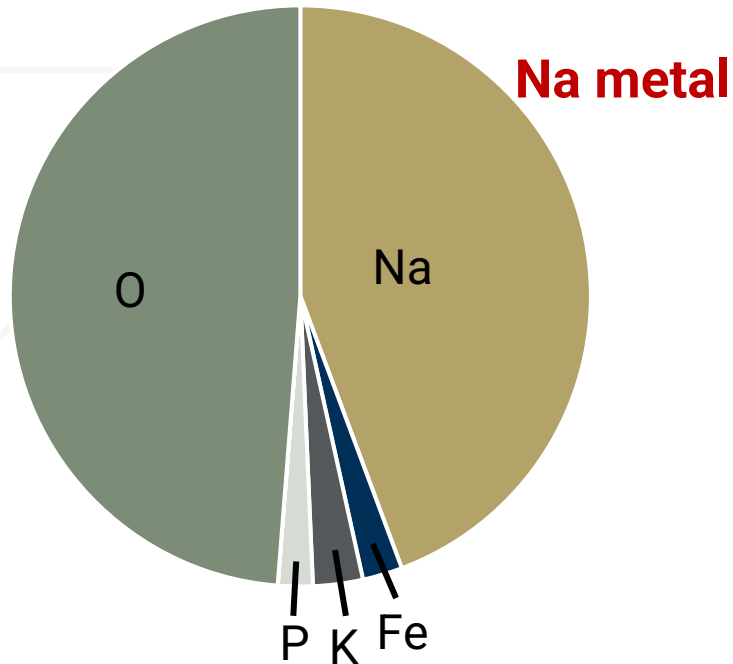


# Direct Thermal Reduction: Material Characterization

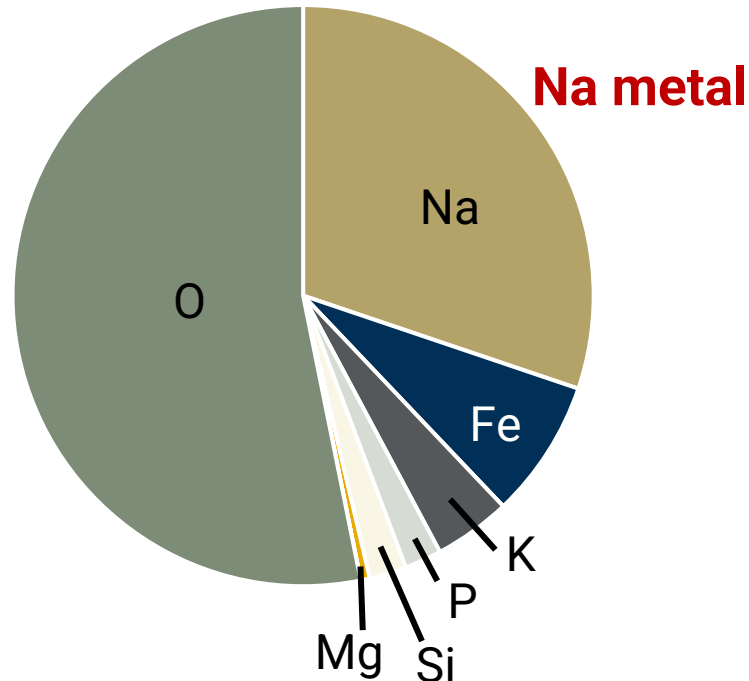


## LMS-1 – XPS

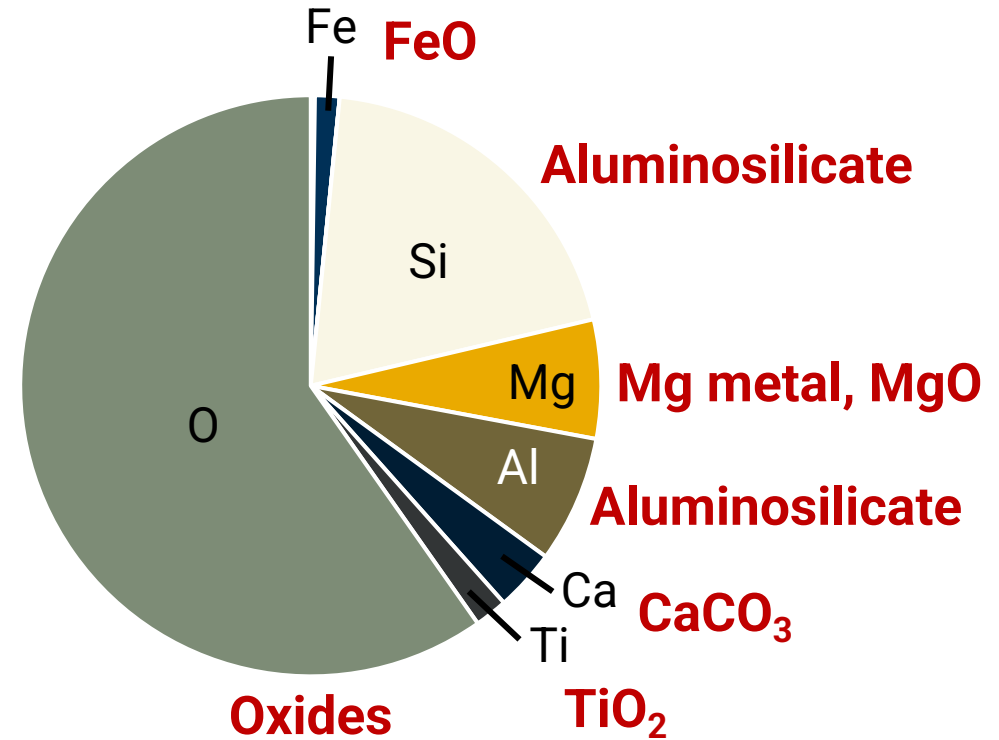
$P = 613.2 \text{ W}$



$P = 730.0 \text{ W}$



Remaining sample



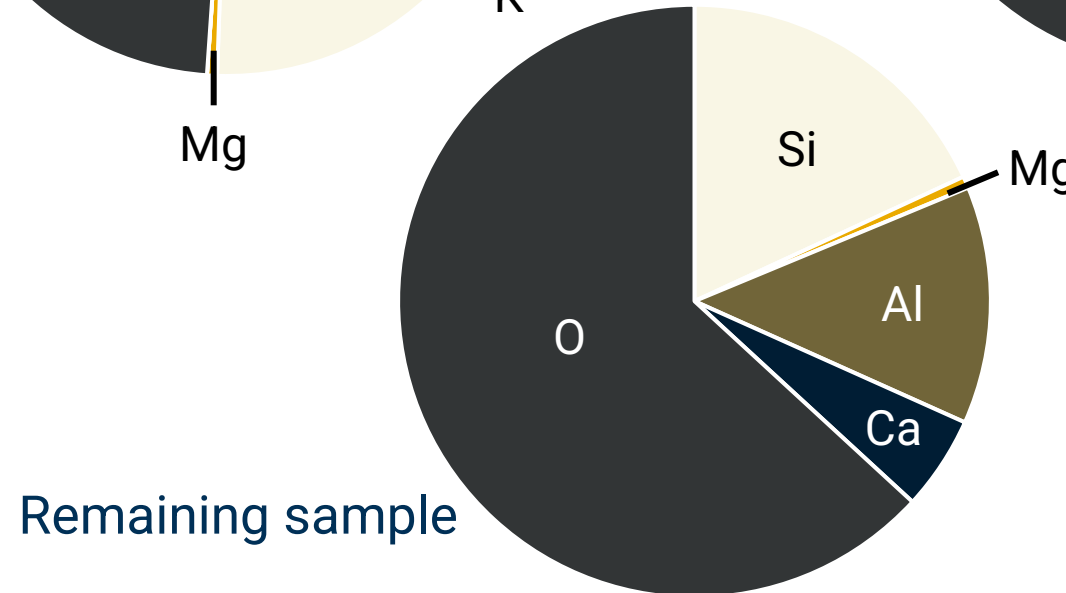
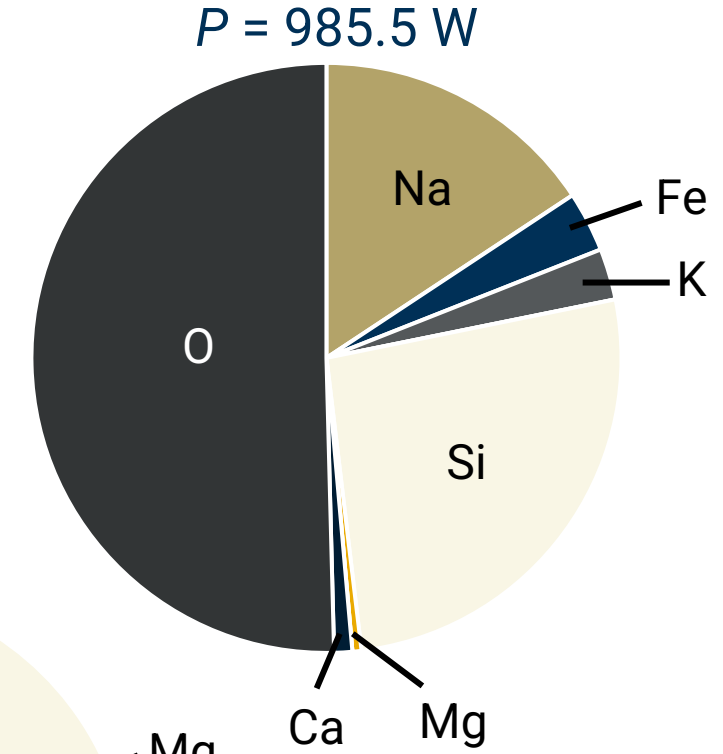
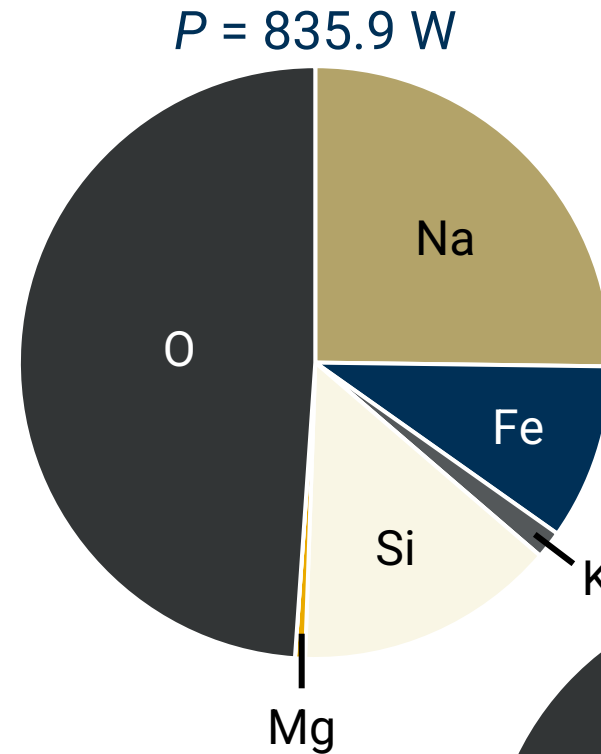
- Depth of analysis for XPS is shallower than EDS

# Direct Thermal Reduction: Material Characterization



## LHS-1 – EDS

- Depositions were evenly distributed on Cu foils
- Na, Fe, and K evolved during both experiments suggesting separation possibility at reasonable temperatures
- Si and Mg evolved continuously, and Al and Ca remained in sample upon experimentation evidencing requirement of elevated temperatures

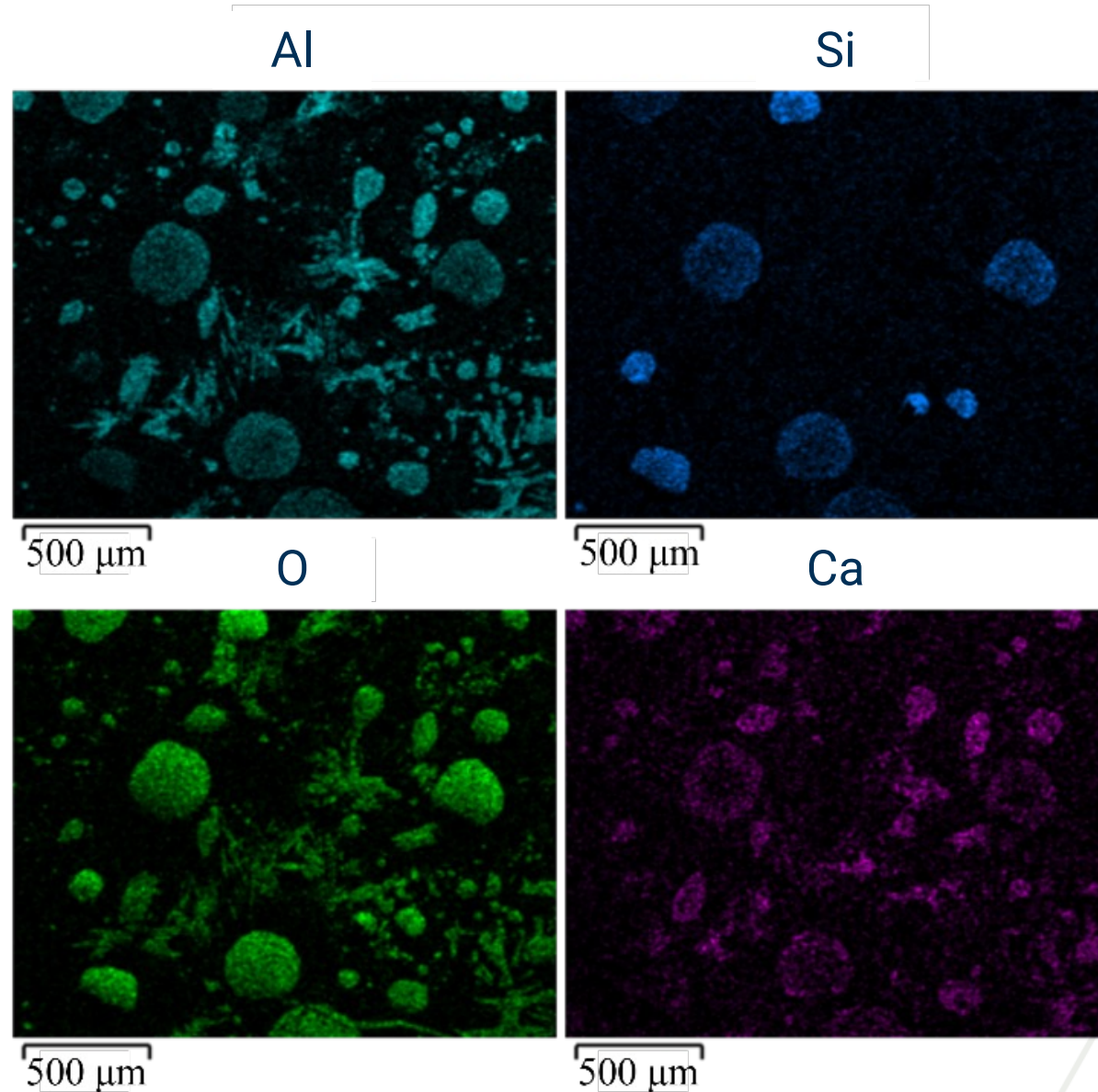


# Direct Thermal Reduction: Material Characterization



## LHS-1 – EDS

- Sample segregation observed on remaining sample
- Significant amounts of sample containing Al and Ca without Si, suggesting potential Al production
- The elements detected are likely oxides, indicated by the O scan

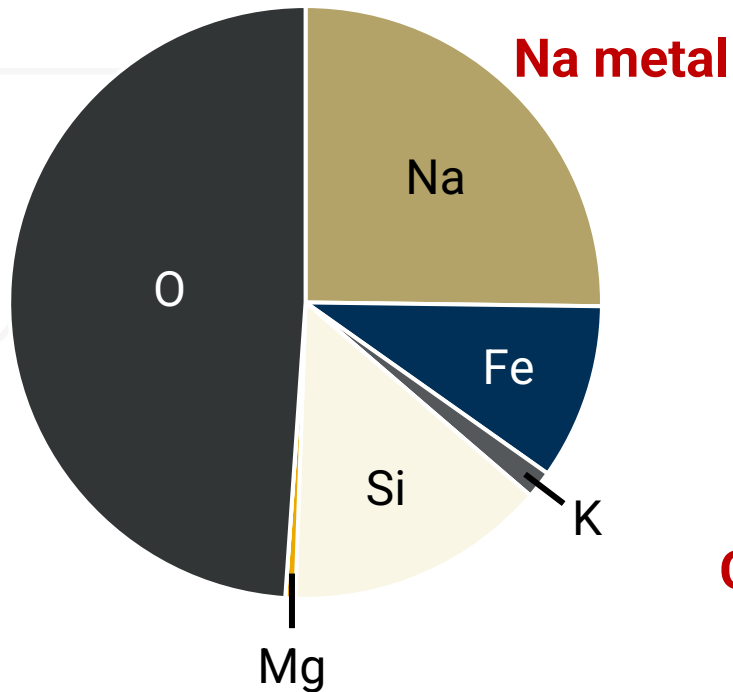


# Direct Thermal Reduction: Material Characterization

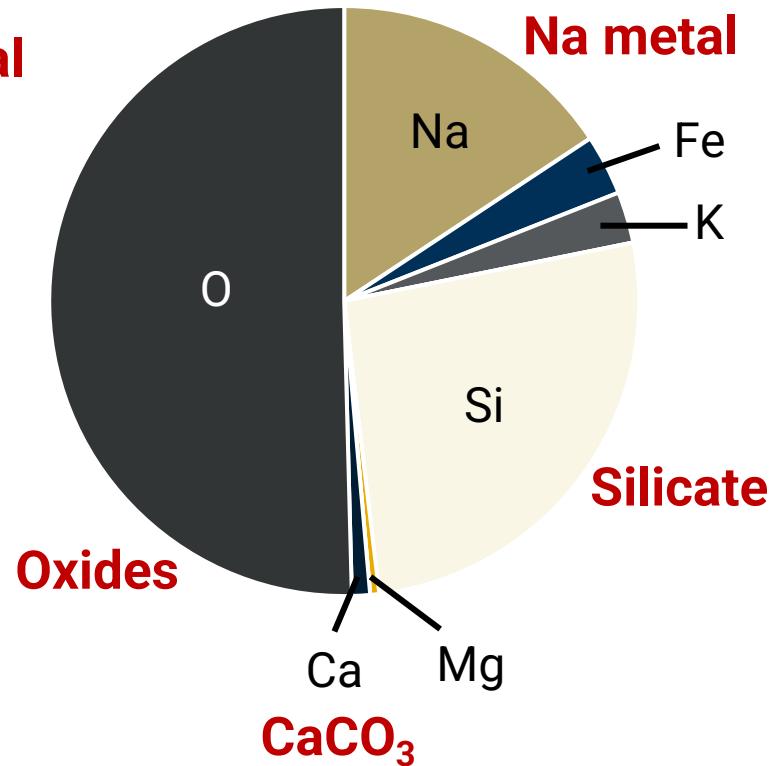


## LHS-1 – XPS

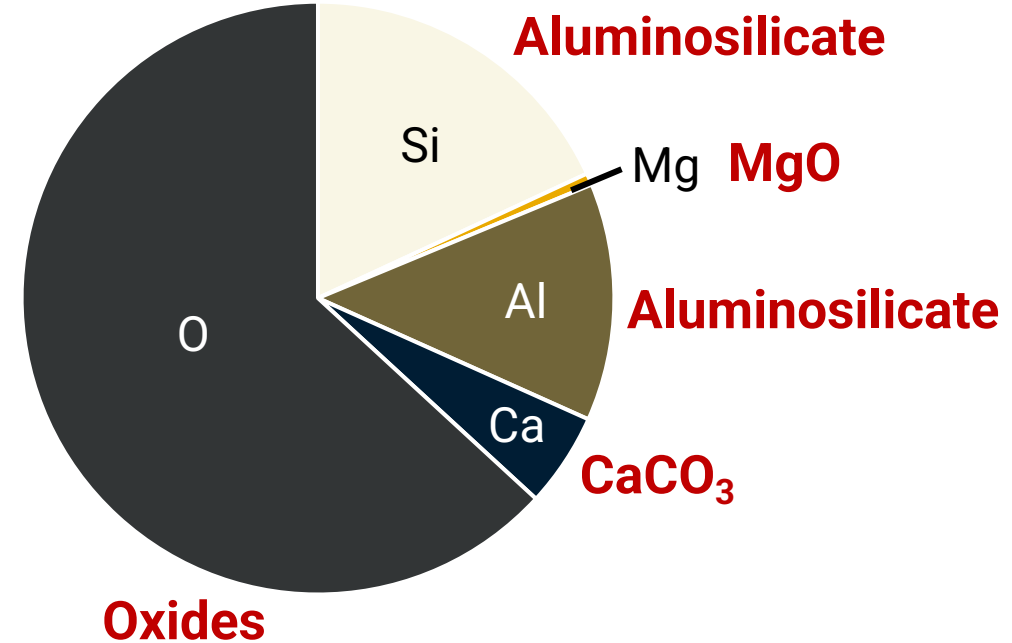
$P = 835.9 \text{ W}$



$P = 985.5 \text{ W}$



Remaining sample

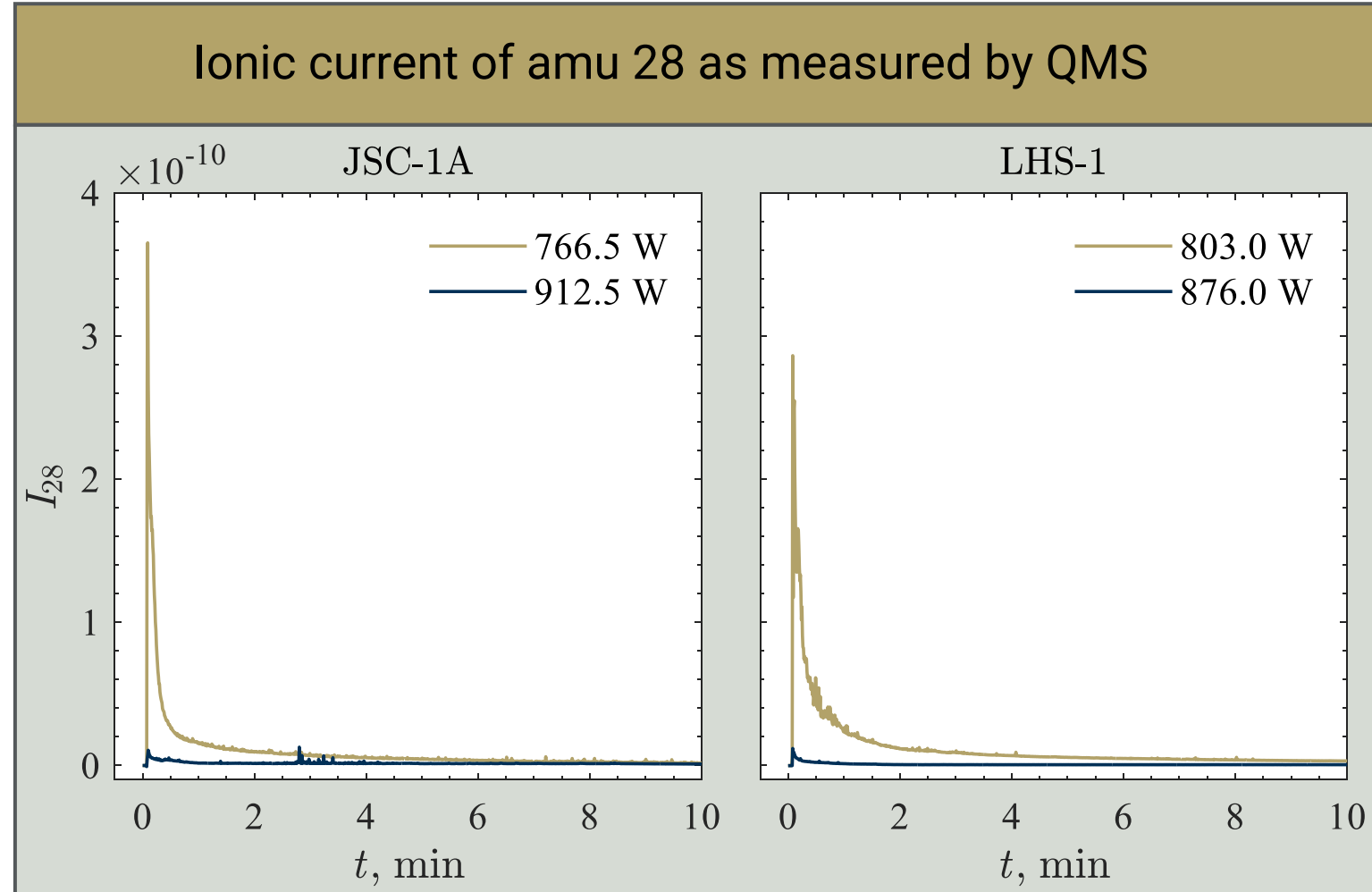


- Depth of analysis for XPS is shallower than EDS



# Carbothermal Reduction

- CO(g) evolution from sample during experiments was evidenced by the QMS at 28 amu
- A prominent peak at  $t \leq 0.5$  min was observed with no significant changes in  $I_{28}$  for  $t > 10$  min, indicative of rapid thermal reductions
- Smaller amounts of CO(g) evolved during the second experiments because the reacted sample from first experiments were used causing smaller sample size

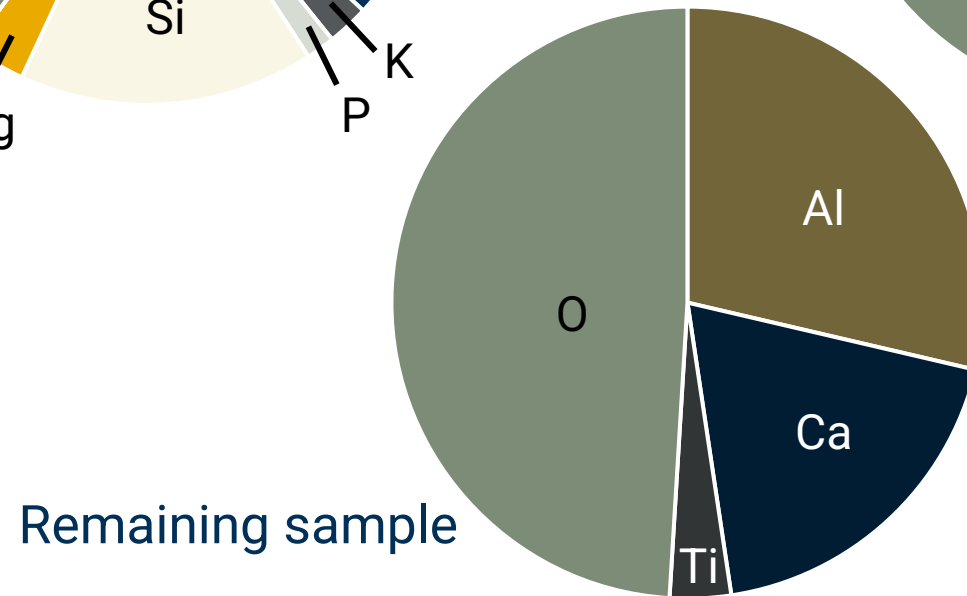
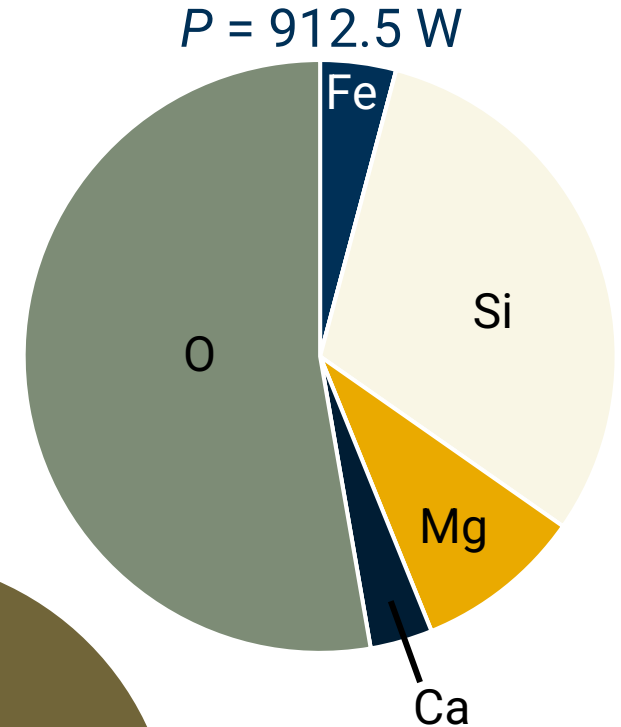
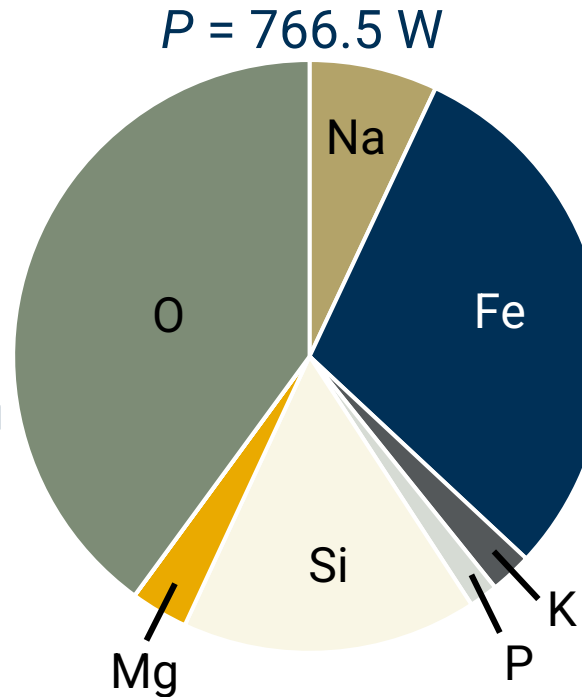


# Carbothermal Reduction: Material Characterization



## JSC-1A + C – EDS

- Depositions were evenly distributed on all foils
- Na, Fe, K, P, Si and Mg evolution during experiments and absence in remaining sample evidenced complete vaporization
- Al, Ca, and Ti remained in sample upon experimentation suggesting separation and production possibility of these elements



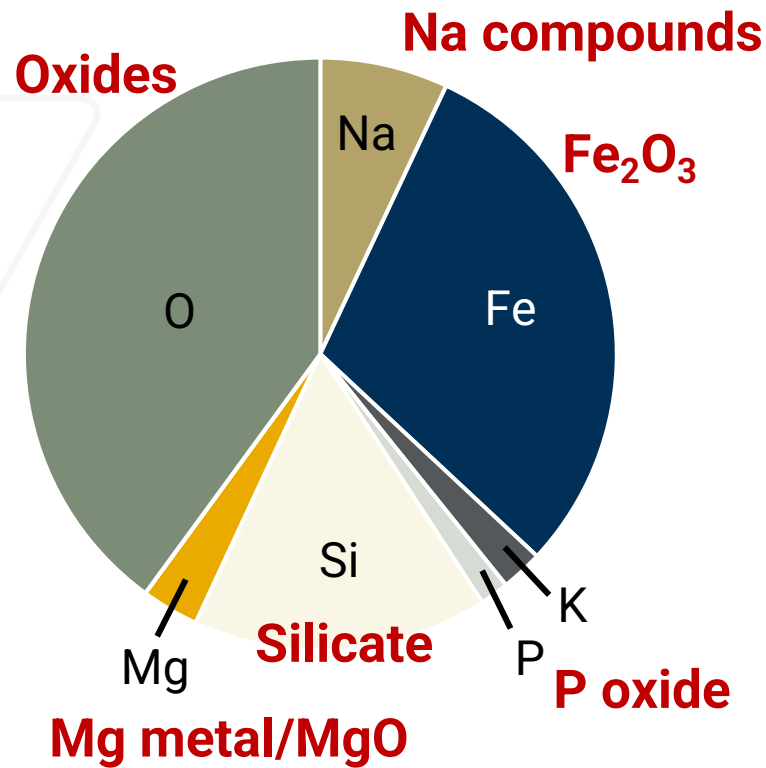


# Carbothermal Reduction: Material Characterization

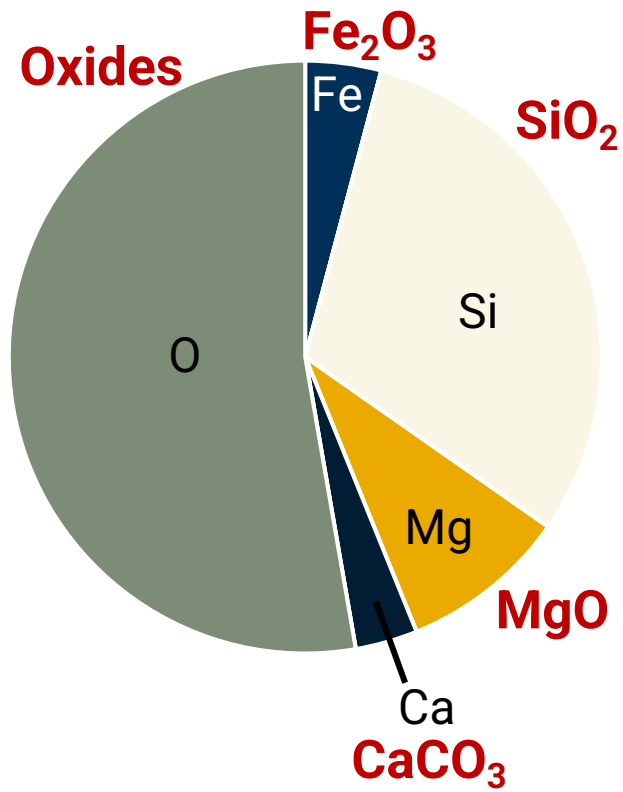


## JSC-1A + C – XPS

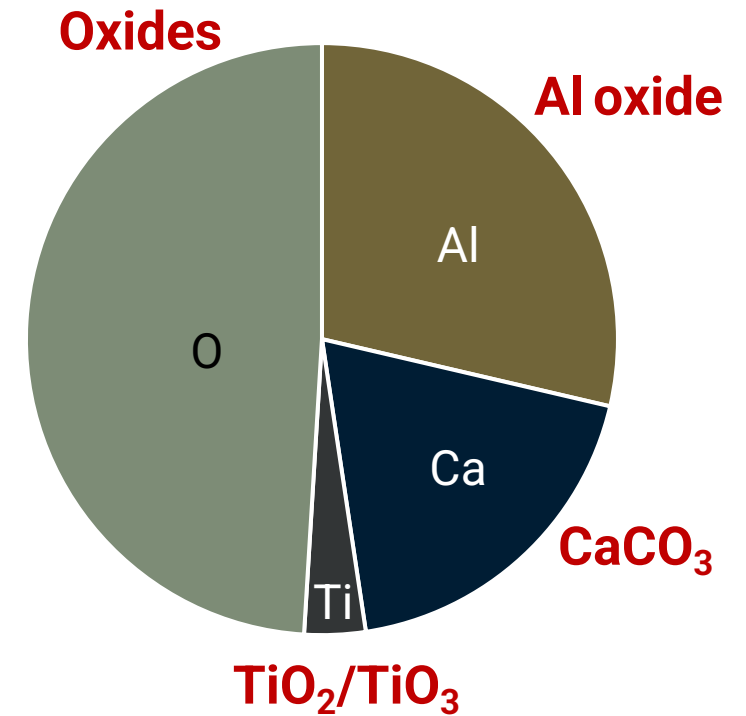
$P = 912.5 \text{ W}$



$P = 766.5 \text{ W}$



Remaining sample



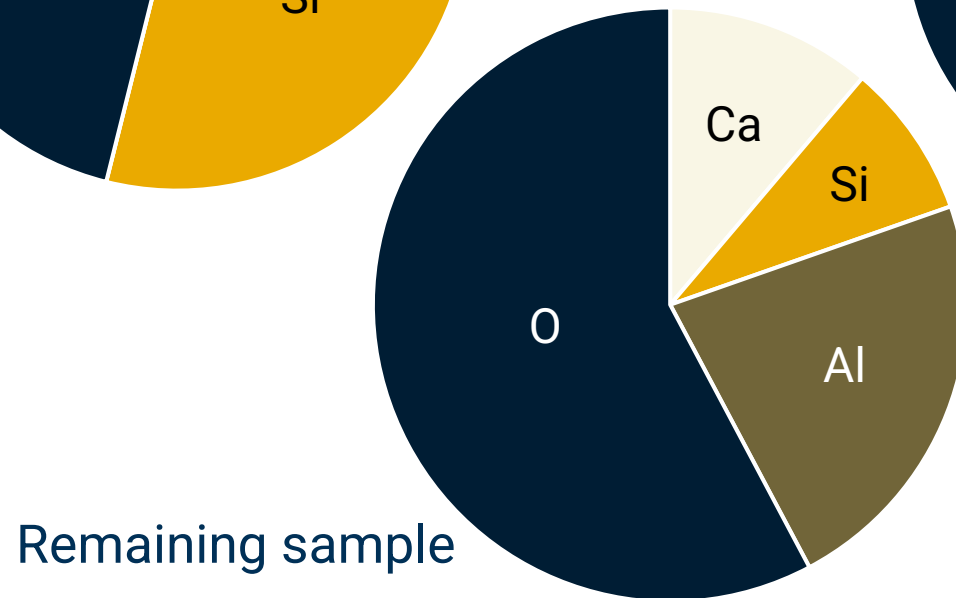
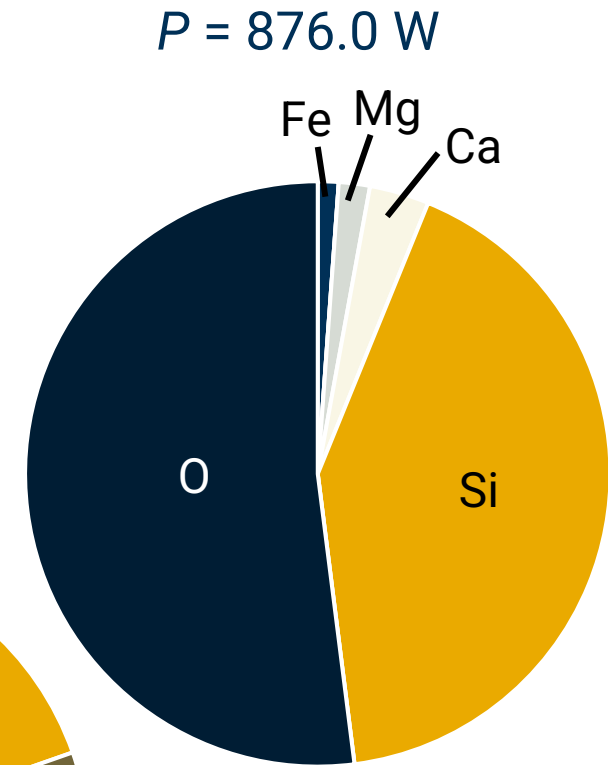
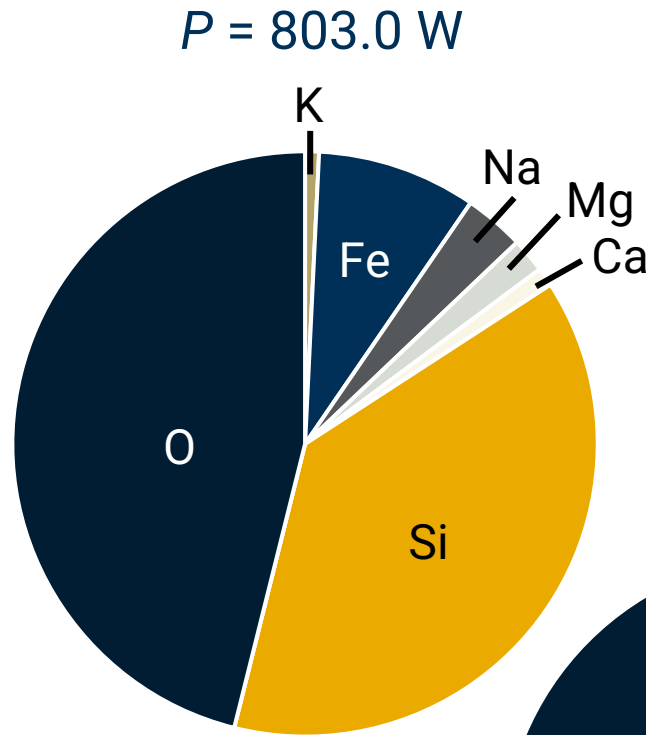
- Depth of analysis for XPS is shallower than EDS

# Carbothermal Reduction: Material Characterization



## LHS-1 + C – EDS

- Depositions were evenly distributed on Cu foils
- Na, Fe, and K evolved during the first experiment suggesting separation possibility at reasonable temperatures and complete vaporization
- Si, Mg, and Ca evolved continuously, and Al, Si, and Ca remained in sample

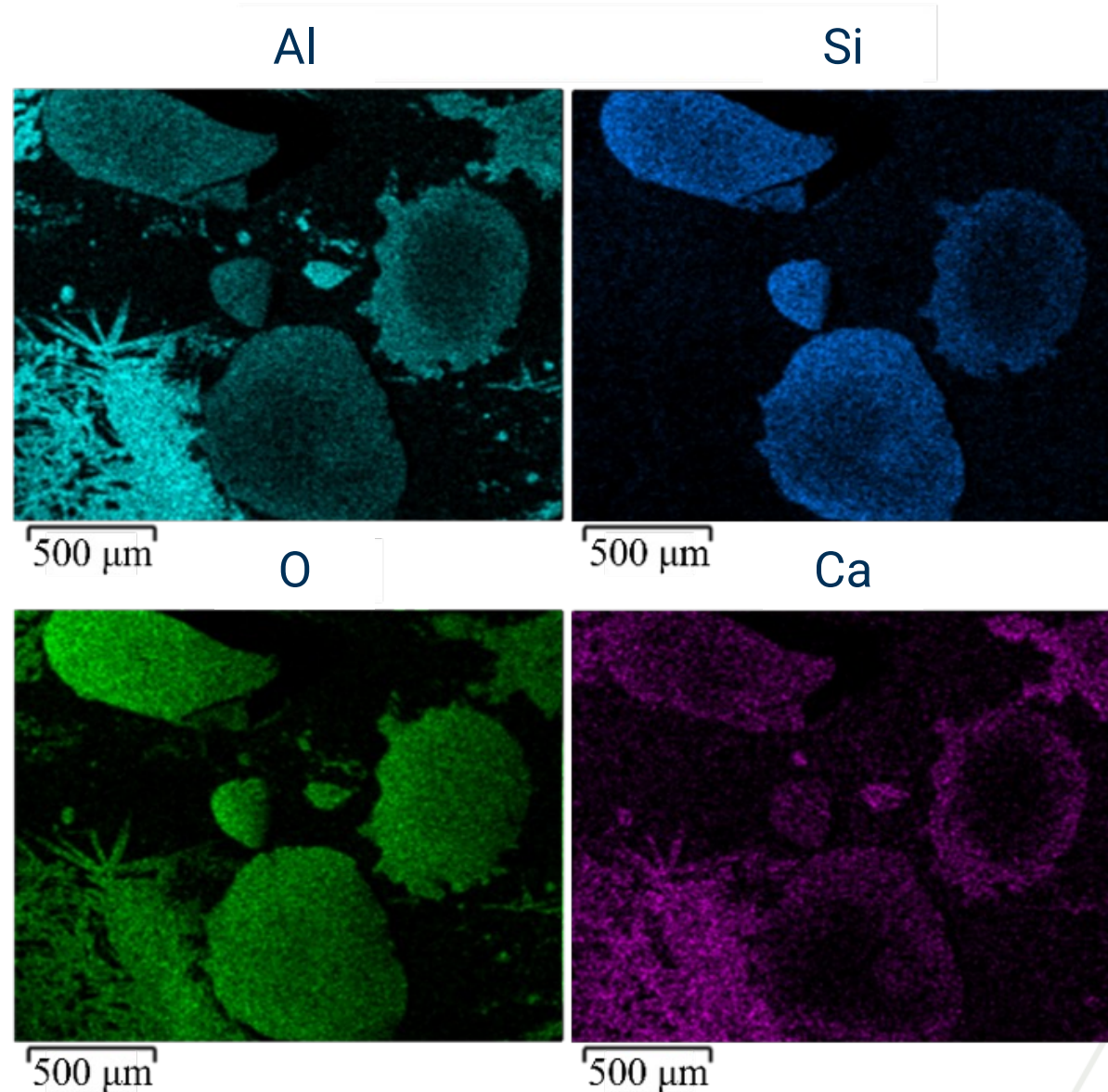


# Carbothermal Reduction: Material Characterization



## LHS-1 + C – EDS

- Sample segregation observed on remaining sample
- Significant amounts of sample containing Al and Ca without Si, suggesting potential Al production
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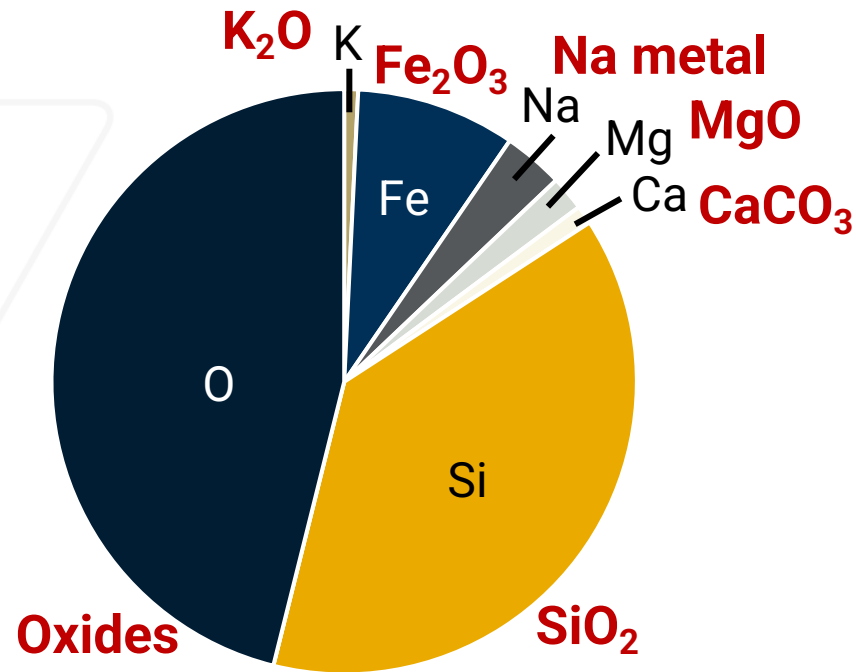


# Direct Thermal Reduction: Material Characterization

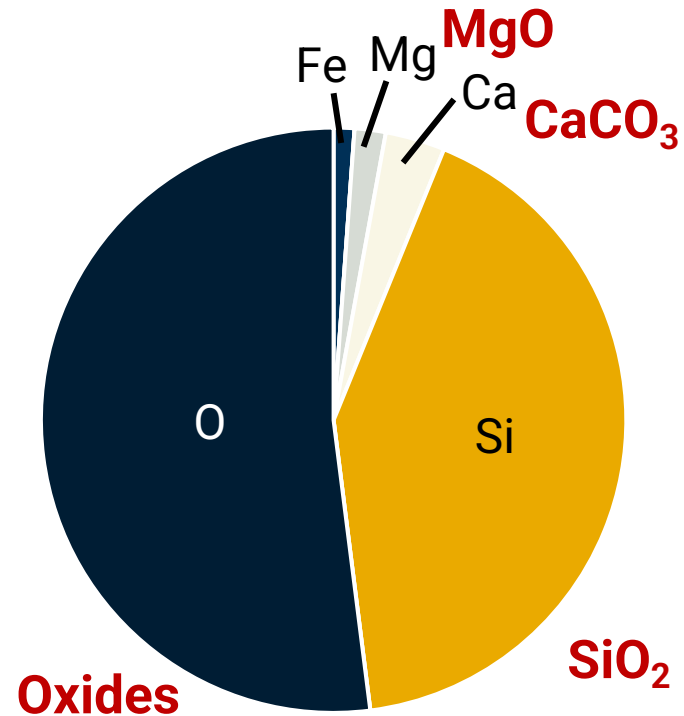


## LHS-1 + C – XPS

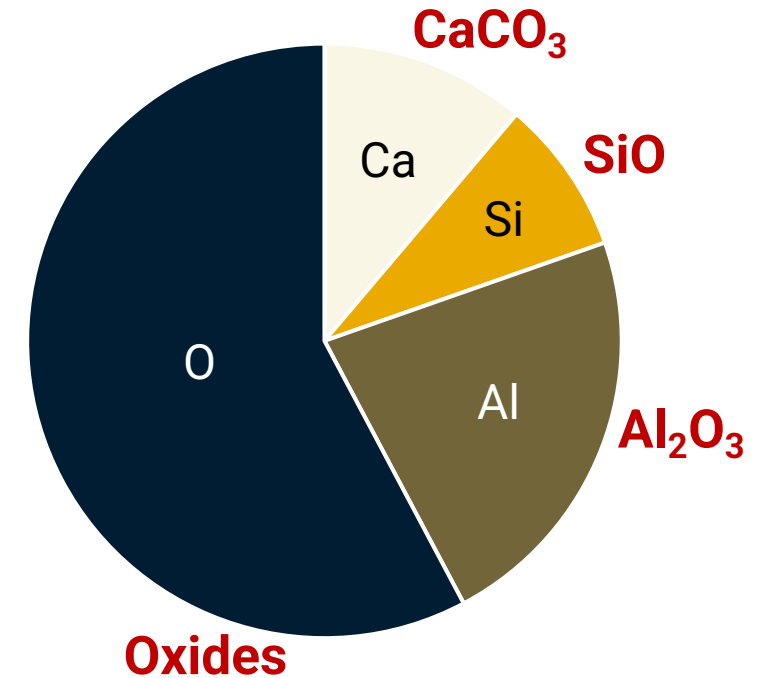
$P = 803.0 \text{ W}$



$P = 876.0 \text{ W}$



Remaining sample



- Depth of analysis for XPS is shallower than EDS

# Summary and Conclusions



- Experimental analyses on thermal and carbothermal reduction of lunar regolith simulants for  $O_2$  and metal/metalloid production in ultra-high vacuum conditions
- Oxygen production was evident in the form of  $O_2(g)$  for thermal reduction and  $CO(g)$  for carbothermal reduction
- Characterization of vapor collecting foils upon experimentation showed that Na, Fe, K, and P are extractable at relatively lower temperatures whereas Si, Mg, and Ca require elevated temperatures
- Characterization of remaining samples upon experimentation showed the possibility of Al separation and production, especially from highlands
- Larger quantities and more compounds vaporized via carbothermal than thermal reduction because reaction onset temperatures were reduced, evidencing the benefit of adding a reductant

# Acknowledgements



## Funders:

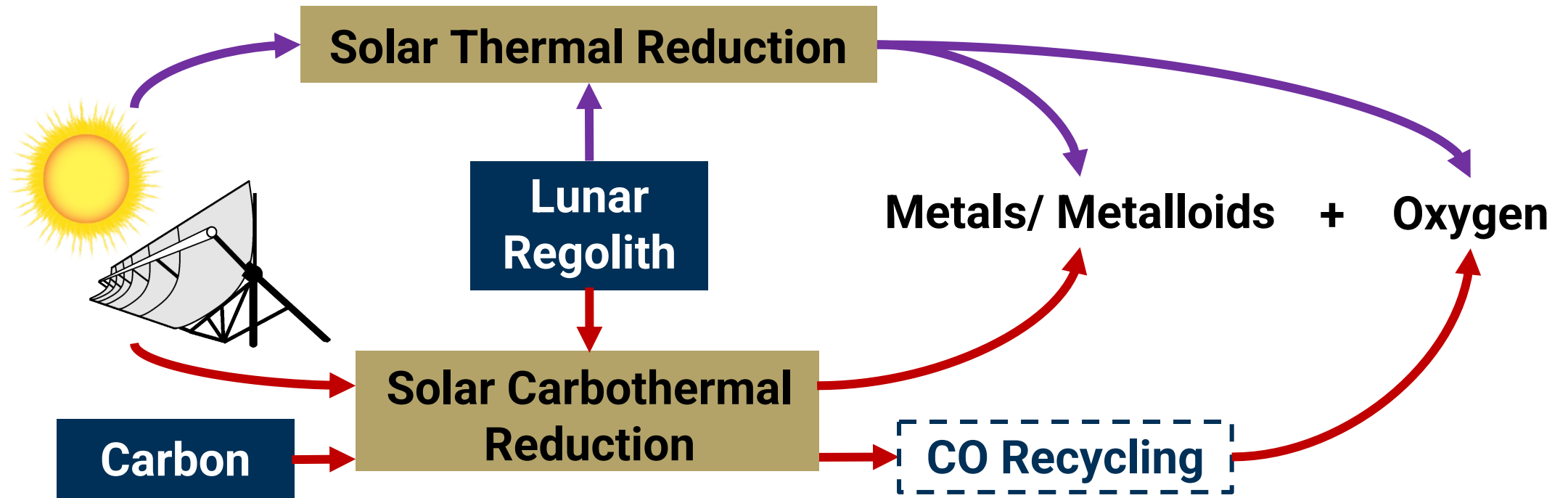
- Lunar *in-situ* Atomic Resource Utilization Experiment (LISA RUE), Georgia Tech Research Institute
- Center for Lunar Environment and Volatile Exploration Research (CLEVER), NASA SSERVI

## Research Team:

- Solar Fuels and Technology Laboratory (Solar FTL) at Georgia Tech
- Electron Proton Induced Chemistry on Surfaces (EPICS) laboratory at Georgia Tech



# Thank you for listening!



## Questions?



# Oxide Reaction Equations

## Direct Thermal

Oxide	Reaction
SiO <sub>2</sub>	SiO <sub>2</sub> → Si + O <sub>2</sub> (g)
Al <sub>2</sub> O <sub>3</sub>	2Al <sub>2</sub> O <sub>3</sub> → 4Al + 3O <sub>2</sub> (g)
FeO	2FeO → 2Fe + O <sub>2</sub> (g)
CaO	2CaO → 2Ca + O <sub>2</sub> (g)
MgO	2MgO → 2Mg + O <sub>2</sub> (g)
Na <sub>2</sub> O	2Na <sub>2</sub> O → 4Na + O <sub>2</sub> (g)
TiO <sub>2</sub>	TiO <sub>2</sub> → Ti + 2O <sub>2</sub> (g)
K <sub>2</sub> O	2K <sub>2</sub> O → 4K + O <sub>2</sub> (g)
P <sub>2</sub> O <sub>5</sub>	2P <sub>2</sub> O <sub>5</sub> → 4P + 5O <sub>2</sub> (g)
MnO	2MnO → 2Mn + O <sub>2</sub> (g)
Cr <sub>2</sub> O <sub>3</sub>	2Cr <sub>2</sub> O <sub>3</sub> → 4Cr + 3O <sub>2</sub> (g)

## Carbothermal

Oxide	Reaction
SiO <sub>2</sub>	SiO <sub>2</sub> + 2C → Si + 2CO(g)
Al <sub>2</sub> O <sub>3</sub>	Al <sub>2</sub> O <sub>3</sub> + 3C → 2Al + 3CO(g)
Fe <sub>2</sub> O <sub>3</sub>	Fe <sub>2</sub> O <sub>3</sub> + 3C → 2Fe + 3CO(g)
CaO	CaO + C → Ca + CO(g)
MgO	MgO + C → Mg + CO(g)
Na <sub>2</sub> O	Na <sub>2</sub> O + C → 2Na + CO(g)
TiO <sub>2</sub>	TiO <sub>2</sub> + 2C → Ti + 2CO(g)
K <sub>2</sub> O	K <sub>2</sub> O + C → 2K + CO(g)
P <sub>2</sub> O <sub>5</sub>	P <sub>2</sub> O <sub>5</sub> + 5C → 2P + 5CO(g)
MnO	MnO + C → Mn + CO(g)
Cr <sub>2</sub> O <sub>3</sub>	Cr <sub>2</sub> O <sub>3</sub> + 3C → 2Cr + 3CO(g)